

KOKURI	Chemical nature of the organic matter in Baltic shale. Trudy IfI  (MIRA 13:8)  no.51:58-64 159.  (Oil shales-Analysis)								
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			January	en e					

# Investigating the acid products of shale oil. Trudy LTI no.51:65-71 159. (Oil shales—Analysis)

### KOKURIN, A.D., GALUTKINA, K.A.

Effect of various admixtures on the yield of acid products in the thermal decomposition of oil shale. Trudy LTI no.51:72-75 159. (MIRA 13:8)

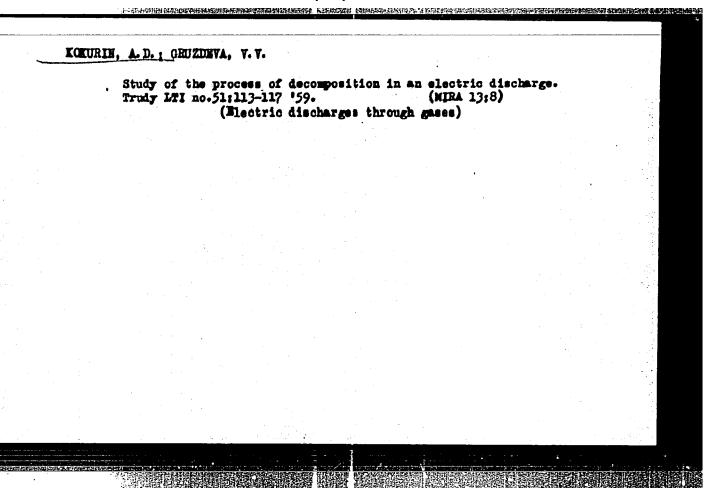
(Oil shales) (Kerogens)

KCKURIN, A.D.; SETKINA, O.N., GRUZDEVA, V.V.

Decomposition of organic matter in an electric arc discharge.

Trudy LTI no.51:102-112 \*59. (NIRA 13:8)

(Hydrocarbone) (Electric arc)



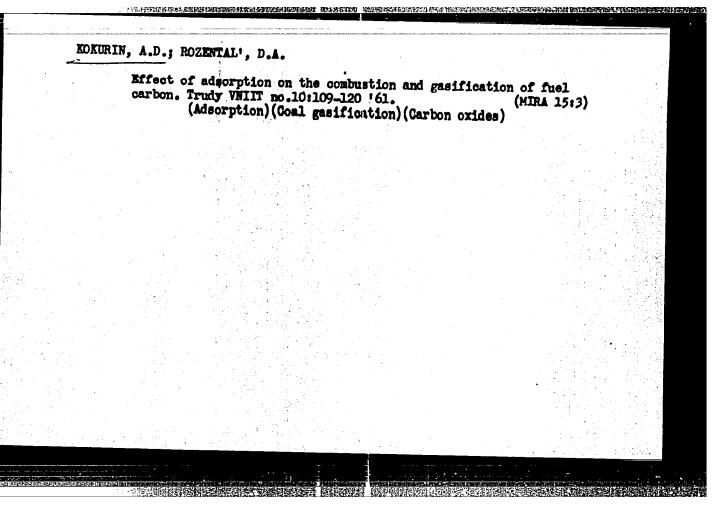
KOKURIN, A.D.; BAL'YAN, Kh.B., nauchny, red.; VOROB'YEV, G.S., red.izd-va; GURDZHIYEVA, A.M., tekhn.red.

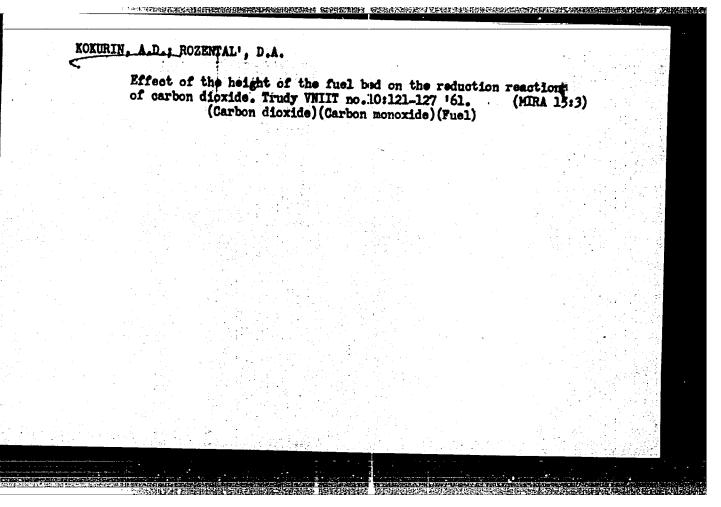
[Role played by chemistry in technological progress] Rol' khimii v tekhnicheskom progresse. Leningrad, Ob-vo po raspr. polit. i nauchm.snanii RSFSR, Leningr.otd-nie, 1961. 52 p.

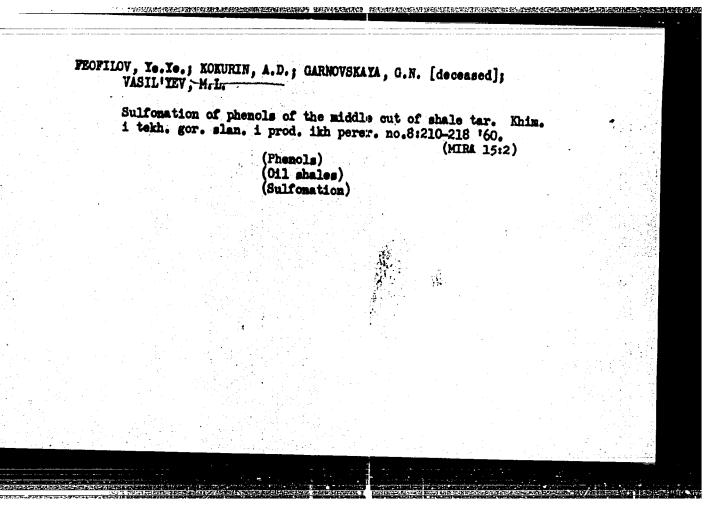
(Chemistry, Technical)

(MIRA 1416)

APPROVED FOR RELEASE: 06/19/2000 CIA-RDP86-00513R000723710017-3"





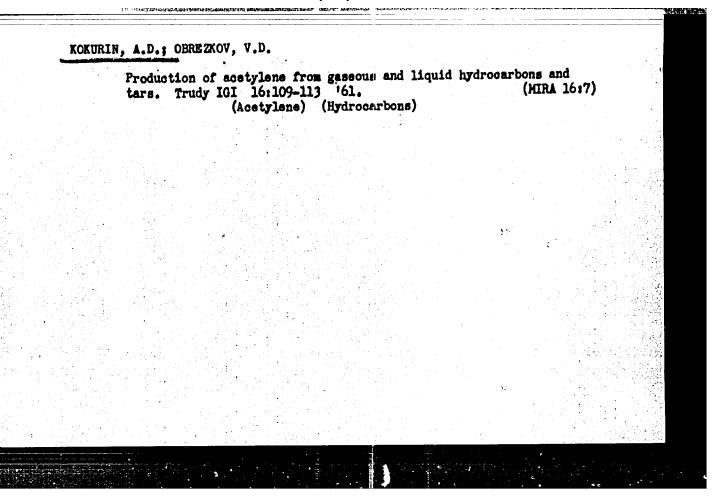


AUTHORS:	S/672/62/000/011/004/011 D403/D307 Kokurin, A. D. and Obrezkov, Y. D.	
TITLE:	Study of the mechanism of the process of electrocracking of liquid products  Leningrad. Veesoyuznyy nauchno-issledovatel'skiy institut pererabotki i ispol'zowaniya topliva. Trudy. no. 11, 1962. Khimiya i takhnologiya topliva i produktov pere-	
arcs, con ing place terials o the crack able elec	brief account is first given of the difficulties of the electrocracking of various liquid products in multiple cluding that little is as yet known of the processes tak- The present study was concluded with simple liquid ma- f known composition. Compositions of gases obtained from ing of benzene with various size fractions of powder mov- trodes (chiefly acetylene (33-37%) and hydrogen (61-66%)), see obtained from toluene and Tylene, are tabulated. The enes gave slightly higher contents of C2H2, C2H4.	

S/672/62/000/011/005/011 D403/D307 AUTHORS: Kokurin, A. D. and Obrezkov, Y. D. The effect of certain factors on the electrocracking in TITLE: microdischarges Leningrad. Vsesoyuznyy nauchno-issledovatel'skiy institut pererabotki i ispol'zovaniya topliva. Trudy. no. 11, SOURCE: 1962. Khimiya i tekimologiya topliva i produktov yego pererabotki. 107-119-TEXT: The object of the present study was to provide guidance for the design of industrial microdischarge electrocracking installations. The work was carried out on laboratory scale. The factors considered were: (1) Construction of apparatus; rectangular reactors are preferable to cylindrical ones. Various reactors are illustrated and compared. The effects of mechanical stirring or shaking of the carbon filling (movable electrodes) are discussed. (2) the effects of the amount and size fraction of the carbon filling on the effectiveness of the process and on the composition of the Card 1/3

S/672/62/000/011/005/011 The effect-of-certain ... resulting gas. In the cracking of kerosene finer filling led to de creased yields of C2H2 and increased C2H4, C3H6 and C4H8. (3) The effects of interelectrode distance (1) and of circulation of the liquid material on effectiveness of the process; the productivity (1/hr) fell sharply with increasing 1, and was slightly higher with circulation. (4) The effect of elemental and group composition of the starting material on the products. The percentage content of OpHo in the gas m is given by  $m = \frac{0.5(n - x)(100 - K)}{n + y}$ where n is the number of Catoms in the molecule, K is the vol. % of C forming soot, and x, y are constants depending on the type of the starting material (for paraffins x = 0, y = 1; for naphthenes x = y = 0; for aldehydes and ketones x = 1, y = 2; for monatomic Abstracter's note: 0,7 alcohols x = 1, y = 2; for diatomic alcohols x = 2, y = 3; for triatomic alcohols x = 3, y = 4). Adjust-Card 2/3

The effect of	certain	S/672/6 D403/D3	2/000/011/005/0 07	11   27
sure (8 - 20 a	omposition is disc ition: reduced pro itm) decreases the	essuro increases of scety	and increased prolessed in the crace	res-
oi kerosene. T	The method of elecate to be highly pr	ctrocracking in m	ultiple voltage	المراديق فوالمتكنية
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Card 3/3				



KOKURIN, A.D.; ROZENTAL', D.A.; YEVDOKIMOV, Yu.P.

Investigating the interaction of oxygen, carbon monoxide and carbon dioxide with charcosl under static conditions. Trudy LTI no.59:101-106 '61. (MIRA 17:9)

KOKURIN, A.D.; ROZENTAL, D.A.; SUSLINA, V.P.; TISHINA, N.I.

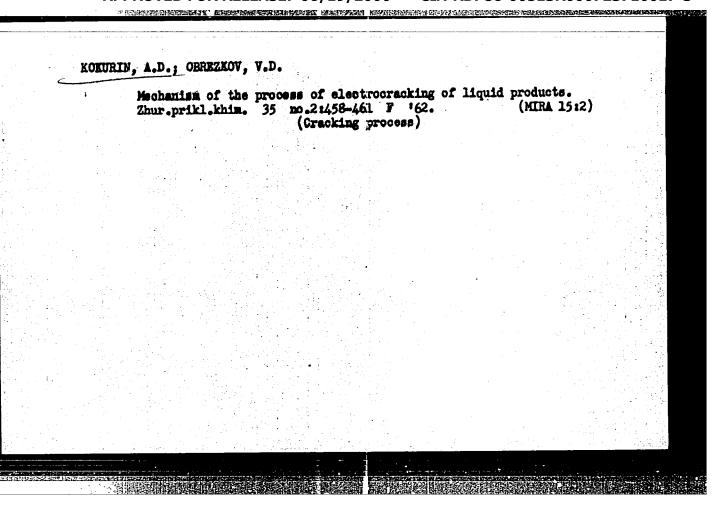
Investigating the interaction of carbon dioxide with fuel carbon under dynamic conditions. Trudy LTI no.59:107-112 '61. (MIRA 17:9)

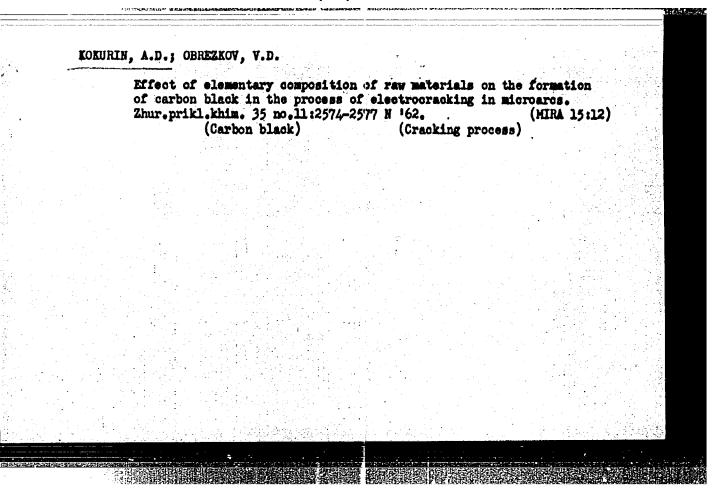
Studying the mechanism of the electrocracking of liquid products.

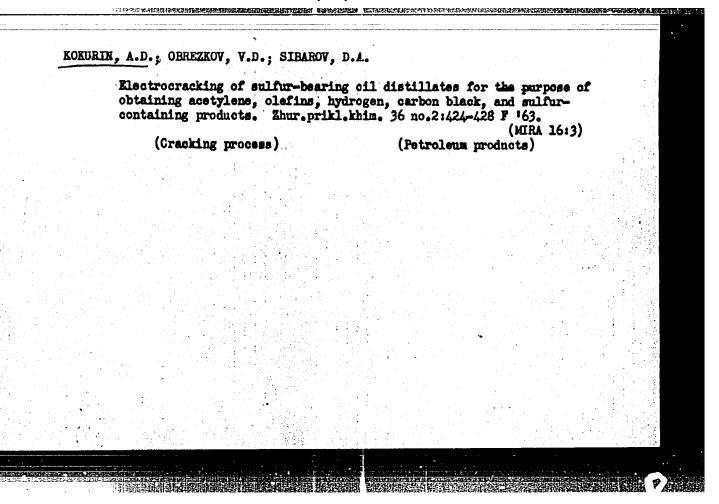
Trudy WNIIT no. 11:101-106 '62.

Effect of certain factors on electrocracking in micro-discharges.

[Did.:107-119. (MIRA 17:5)





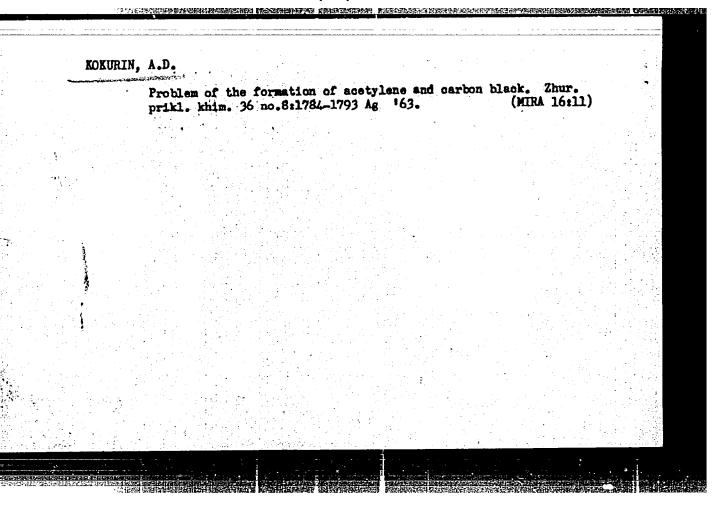


KOKURIN, A.D.; OBREZKOV, V.D.; ANDREYEV, N.S.

Preparation of vinyl acetate from diluted acetylene. Zhur. prikl. (MIRA 16:7)

(Vinyl acetate) (Acetylene)

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### "APPROVED FOR RELEASE: 06/19/2000

CIA-RDP86-00513R000723710017-3

ACC NR. AP6025613

(A)

SOURCE: CODE: UR/0413/66/000/013/0053/0053

INVENTOR: Sibarov, D. A.; Kokurin, A. D.; Krzhechkovskiy, G. N.

ORG: None

TITLE: A device for studying electric discharges in liquids. Class 23, No. 183312

SOURCE: Izobreteniya, promyshlennyye obraztsy, tovarnyye znaki, no. 13, 1966, 53

TOPIC TAGS: electric discharge, electrode

ABSTRACT: This Author's Certificate introduces a device for studying electric discharges between stationary and movable electrodes in liquids. Isolated electric discharges are produced by mounting the stationary electrode on the bottom of the vessel for the liquid with the movable electrode suspended above it on a flexible lead.

BUB CODE: 09, 20/ BUBM DATE: 17May65

Card 1/1

WC; 66,092,193,05

POLOZOV, V.F.; ZAPEVALOV, N.V.; SOTNIKOV, M.A.; KOLODIN, E.A.; KOKURIN, A.D.

Breaking down kerosine in mcmantary intermittent electric arcs.

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and the state of t			<b>上的作用框</b>
L 56533-65 ACCESSION NR: AP5016790	UR/C	286/65/000/010/0138	/0138
AUTHOR: Dobryanskiy, A. F.; Ko	66.0	97.3 : 691.4	Z Tarana
TITLE: A method for producing SOURCE: Byulleten' izobreteniy	. is all		
TOPIC TAGS: active medium, cla ABSTRACT: This Author's Certifi by acid treatment. The clay is	Flante introduces a mutho	i for producing acti	ve clay 10-17 at-
mospheres in water with the add treated in acid.	ittion of catalysts,g.	mineral acid salts	and then
ASSOCIATION: none SUBMITTED: 16Dec50	ENCL: CO	SUB CODE: ES	, <b>cc</b>
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NIKITIN, Ye. A.;	KOKURIN, A.				
Acids, Inorganic			1	1050	
Structure of hete	eropolyacids,	Izv. Sekt. plat. i bla	. met., No. 25,	1950.	
				3	, Uncl.
9. Monthly L	ist of Russian	Accessions, Library o	f Congress, Apr	1958	, oner-
		Mariana A.			

KOKUKIN

SUBJECT:

USSR/Schooling

27-8-8/32

AUTHOR:

Kokurin, I.

TITLE:

Systematic Repair of Equipment (Planovost' v remonte oborudo-

vaniya)

PERIODICAL:

Professional'no - Tekhnicheskoye Obrazovaniye, Aug 1957, #8,

p 14 (USSR)

ABSTRACT:

Until 1956, insufficient attention was paid in Tula to the preventive maintenance of trade school equipment. This condition improved after the Chief Administration of Labor Reserves, on 21 April 1956, issued an order concerning the organization of systematic preventive maintenance. As a result, 21 schools which lacked the necessary equipment have established at the Special Trade School # 30 a community, well-equipped machine shop to take care of the required maintenance work. The other city trade schools are mechanically well equipped to care for their own needs. Further improvement, however, is still required.

INSTITUTION: Tul'skoys Oblastnoys Upravleniye Trudovykh Rezervov (The Tula Provincial Administration of Labor Reserves).

PRESENTED BY:

SUBMITTED:

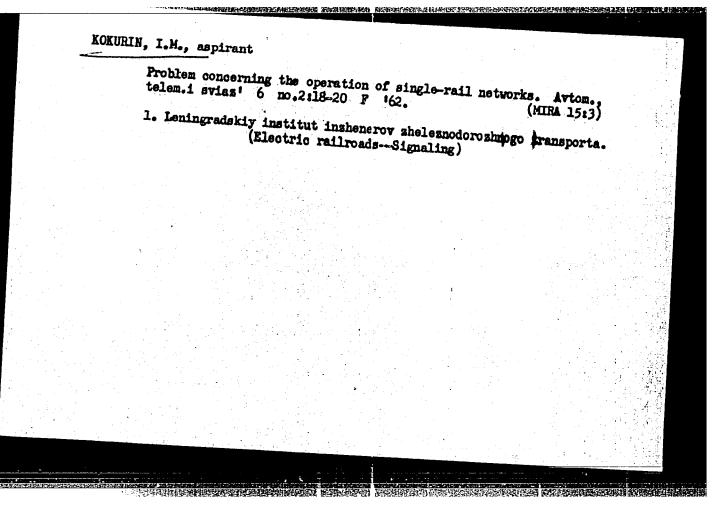
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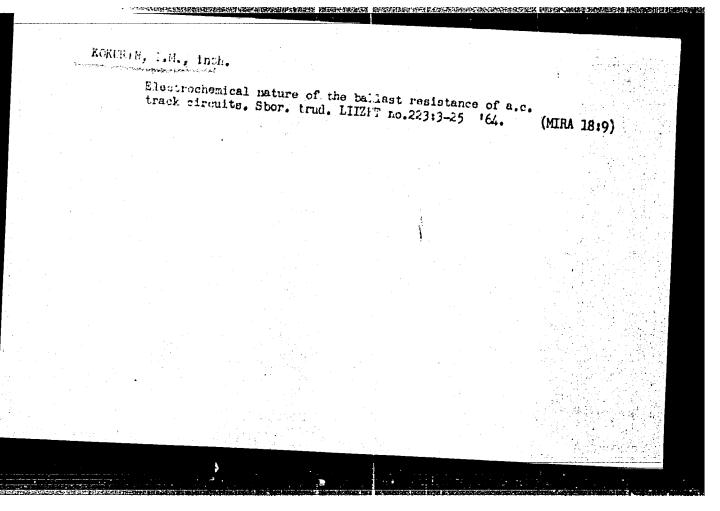
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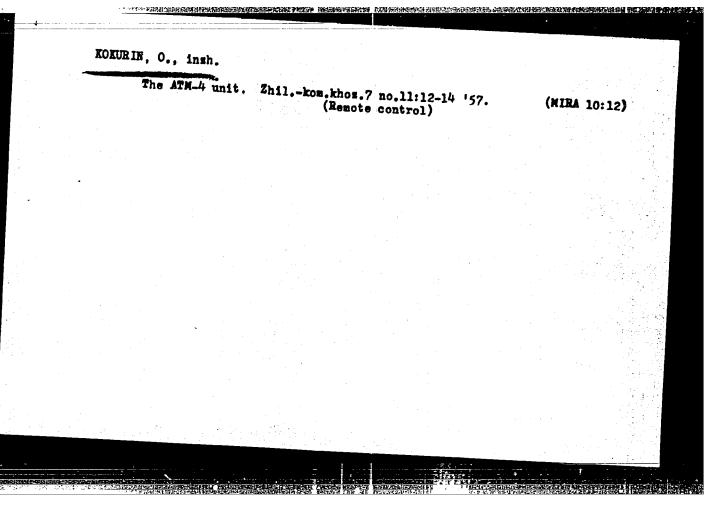
Card 1/1

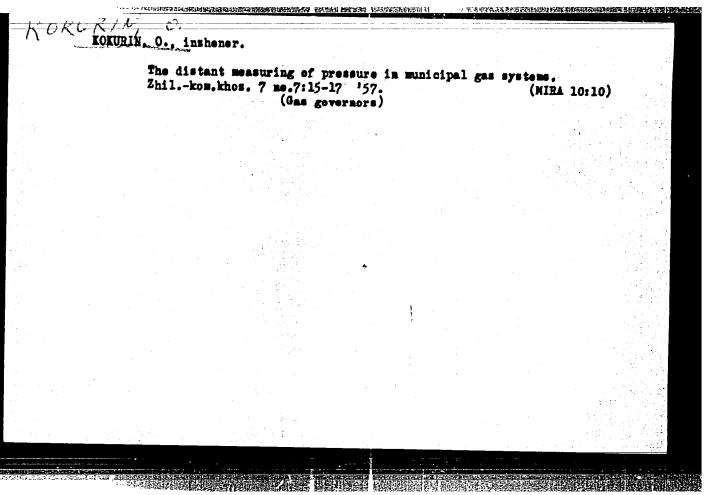
CIA-RDP86-00513R000723710017-3" APPROVED FOR RELEASE: 06/19/2000

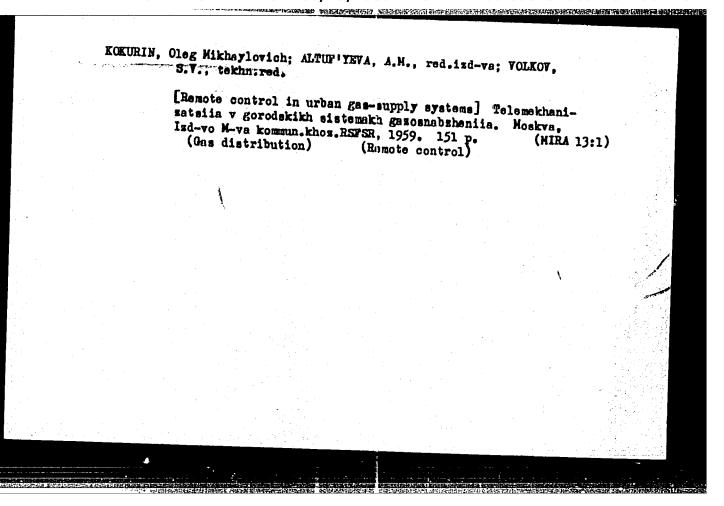
计类型点 化油气蒸馏剂油罐 6(4):28(1) PHASE I BOOK EXPLOITATION Cont. T. M. Partingue, V. C. Kie-Kokurin, Ivan Ivanovich, and Naum Samoylovich Payngersh Avtomatizatsiya upravleniya radiouslami (Automation of Rediffusion Stations) Moscov, Svyaz izdat, 1958. 53 p. (Series: Ogyt peredovykh svyazistov) Resp. Ed.: I. P. Bushin; Ed.: L. I. Vengrenyuk; Tech. Ed.: K. G. Markoch. PURPOSE: The booklet is intended for specialists in rediffusion broadcasting. COVERAGE: The authors, both specialists in the automation of repeater stations in rediffusion broadcasting, describe in detail the methods used in converting amplifying systems to remote-controlled operation in the town of Gor'kiy. They describe experience gained in operating the equipment, in checking its performance and in telemetering and remote monitoring of conditions in transmission lines. They also list the advantages in economy resulting from the automation of rediffusion stations and substations. The authors conclude that experience gained in the automation of the broadcasting network in the town of Gor'kiy can be used in other cities. The following persons participated











BERSENEV, Ivan Sergeyevich; KOKURIN, O.M., red.

[Automatic and remote control in supplying gas to cities]
Avtomatika i telemekhanika v gazosnabahenii gorodov. Mc-skva, Stroiizdat, 1964. 170 p. (MIRA 18,3)

。 一个大学,我们就是一个大学的,我们就是一个大学的,我们就是一个大学的,我们就是一个大学的,我们就是一个大学的,我们就是一个大学的,我们就是一个大学的,我们就是一个

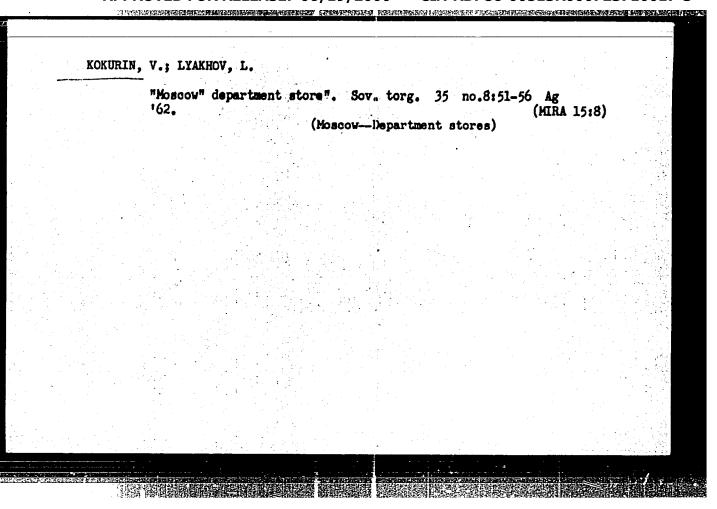
KOKURIN, V.

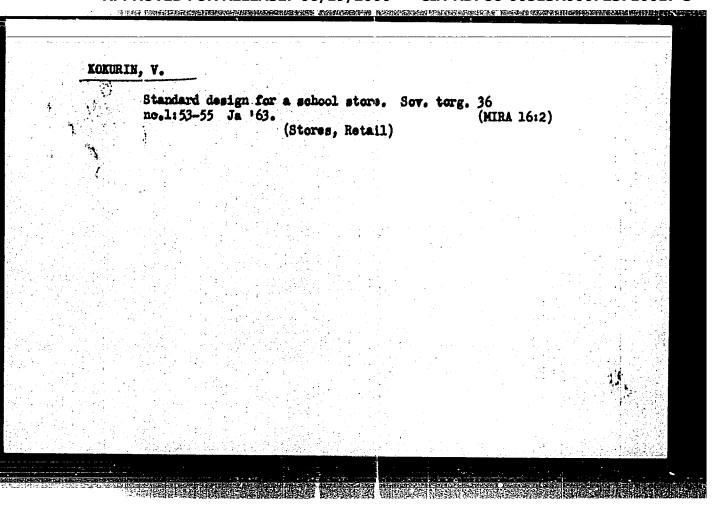
Some problems of planning and designing. Obshchestv.pit. no.8:36-39 Ag '59. (MIRA 12:12)

1. Upravlyayushchiy Moskovskim filialom Giprotorga. (Restaurants, lunchrooms, etc.)

APPROVED FOR RELEASE: 06/19/2000 CIA-RDP86-00513R000723710017-3"

	KOKURIN,	٧.									•
•	-	New Hy	pickling '62,	and	packaging	shops.	Sov.	torg.	35	no.5:59-61 (MIRA 15:5)	
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Class, ceramics, plastics. Obshchestv.pit. no.2:53-54 F \*63.

(MIRA 16:4)

1. Direktor Gosudarstvennogo instituta po proyektirovaniyu predpriyatiy torgovli i obshchestvennogo pitaniya (for Kikurin).

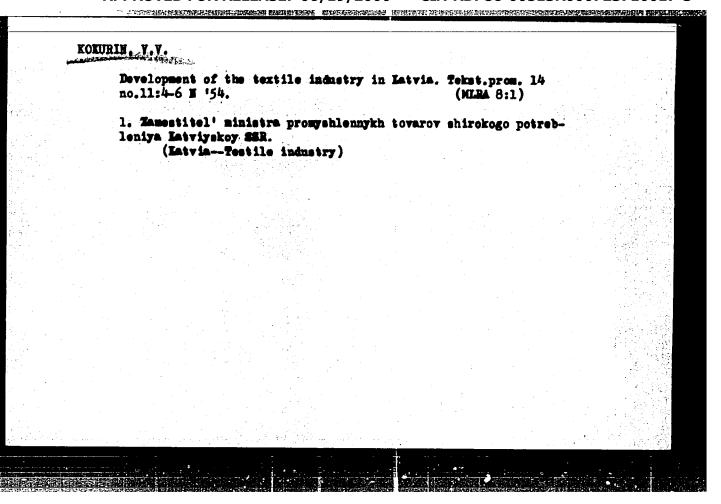
(Restaurants, lunchrocms, etc.—Design and construction)

KOKURIN, Vyacheslav Aleksandrovich; SOFRONOV, V., gornyy inzh.,
retsenzent; SHIRMOV, A.A., gornyy inzh., retsenzent;
KOLOMITTSEV, A.D., gorn. inzh., otv. red.; LYUBIMOV, N.G.,
red.izd-va; IL'INSKAYA, G.M., tekhn. red.; LOMILINA, L.N.,
tekhn. red.

[Assistant engineer of electric locomotives in open-cut mines]
Pomoshchnik mashinista elektrovoza na kar'erakh. Moskva, Gosgortekhizdat, 1963. 282 p. (MIRA 16:12)

1. Nachal'nik vmutrikar'yernogo transporta tresta "Korkimugol'"
(for Sofronov).
(Mine railroads)

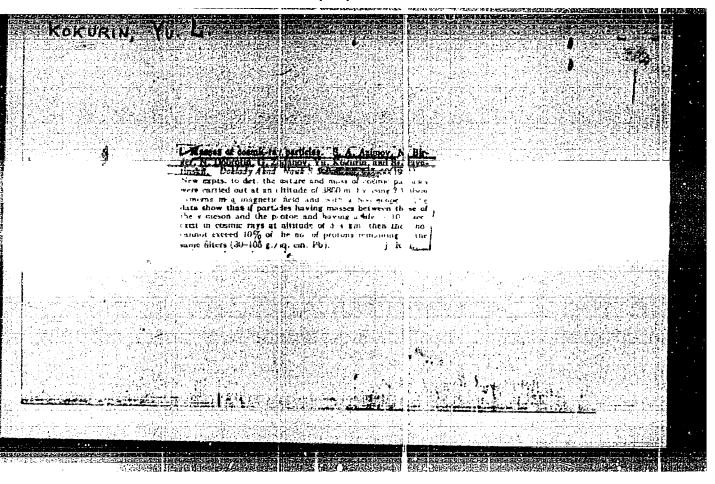
("Slactric locomotives)

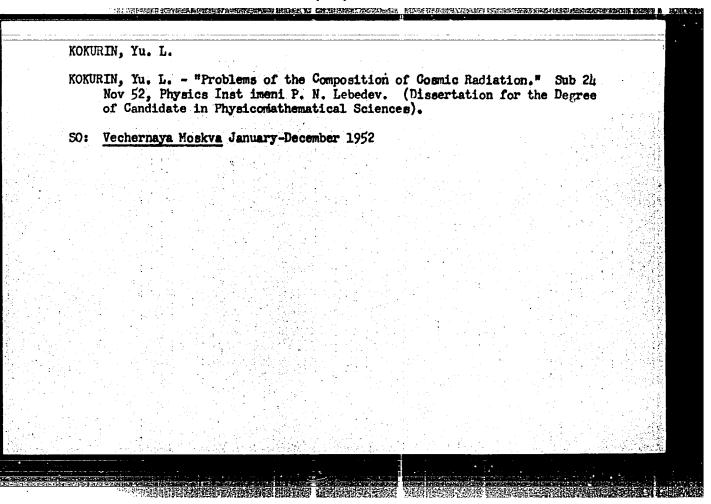


39484-66 EWT(d)/EWP(1) IJP(c) GG/ ACC NR: AT6002989 SOURCE	CODE: UR/0000/6	5/000/000/0213	/0219
UTHOR: Kokurin V.A			
RG: none	riika ka k		
ITLE: Frequency-to-binary-and-decim	al-number converte		
OURCE: <u>Ysesoyumnoye soveshchaniye</u> p			
ychislitel'noy tekhniki. 9th, Yerevan, Magnetic digital elements); doklady sove 13-219	1963. Magnitnyye ti	sifrovyye elemen	ity 🤼
OPIC TAGS: transducer, converter, fr	ediency converter.	frequency code	
onverter, frequency meter			
BSTRACT: A <u>frequency-to-number co</u> n	werter designed on	the electronic-co	ounter 🕍 🤾
rinciple is intended for delivering (deci- pinary) information to a digital computer			
onverter, where fr is the reference-geneasurand (500 cps-5 kc). The convert	nerator frequency (	kc) and f_ is t	he di
Cord 1/2	er operates with ave	raging times of	230, 120

# KOKORIN, K.V.; KOKURIN, V.V.; MEDVEDEV, V.I. Ways to achieve a further upswing of the textile industry. Tekst. prom. 22 no.8:5-8 ag '62. (MURA 15:8) 1. Zamestitel' predsedatelya Ivanovakogo soveta narodnogo khosyaystva (for Kokorin). 2. Machal'nik proisvodstvenno-tekhnicheskogo otdela Ivanovskogo soveta narodnogo khosyaystva (for Kokurin). 3. Zamestitel' nachal'nika Tsentral'nago byuro tekhnicheskoy informateii Ivanovskogo soveta narodnogo khosyaystva (for Medvedev). (Textile industry)

## KOKURIN, V.V.; MEDVEDEY, V.I. Patriotic initiative of assistant foreman.M.V.Kulikova in operation. Tekst.prom. no.2:1-3 7 '63. 1. Nachal'nik proisvodstvenno-tekhnichekkogo otdela Ivanovskogo soveta narodnogo khosyaystva (for Kokurin). 2. Zamestitel' nachal'nika TSentral'nogo byuro tekhnicheskoy informatsii Ivanovskogo soveta narodnogo khosyaystva (for Medvedey). (Textile industry) (Efficiency, Industrial)





VILSOF, John Grahas, 1911; ROBURIS, Yu.L.[translator]; BARADEST, L.T.
[translator]; LABSKIY, L.G., Emidomini; CHERRISTOV, P.A., redaktor;
GERASIMOVA, R.S., tekinioheskiy redaktor.

[The principles of cloud-chamber tekinique; translated from the
English] Kamera Vil'sona. Perevod s'angliiskogo IU.L. Kokurina i
L.T. Baradsei. Moskva, Ind-vo inostramnoi lit-ry, 1954. 151 p.

(Gloud chamber)

(MIRA 7:8)

"Research on the Irregular Structure of the Ionosphere by means of Radioastronomical Methods,"

paper presented at 12th General Assembly of the International Scientific Radio Union [URSI] at Boulder, Colorado, 22 Aug - 5 Sept 57

**どした ロボノ** 

AUTHOR: TITLE:

VITKEVICH, V.V., KOKURIN, YU.L.

109-7-2/17

Radiowave Refraction Irregularities and Considerable Discontinuities in the Ionosphere. (Neregulyarmosti refraktsii radiovoln

i bol'shiye neodnorodnosti v ionosfere, Russian)

Radiotekhnika i Elektronika, 1957, Vol 2, Nr 7, pp 826-832

(U.S.S.R.)

ABSTRACT:

PERIODICAL:

The measuring method and the results obtained by the investigation of the vertical refraction of radio waves by cosmic sources in zenith angles of s = 90 - 70°, at a wave length of 4 m is described. It is shown that vertical refraction is frequently subject to irregular modifications and that irregular refraction is a result of the occurrence of electron heterogeneities with horizontal measurements of the order of 200 km in a height of ~ 350 km (F-layer). At such a distance the optical layer thickness can change by 15-20%.

Several models of the heterogeneous ionosphere are investigated. The deily development of the heterogeneities is analyzed, and it is shown that the occurrence of heterogeneities in the ionosphere is connected with sun activity. (With 7 Illustrations,

1 Table and 1 Slavic Reference).

Card 1/2

APPROVED FOR RELEASE: 06/19/2000 CIA-RDP86-00513R000723710017-

> 109-7-2/17 Radiowave Refraction Irregularities and Considerable Discontinuities in the Ionosphere.

ASSOCIATION:

Physical Institute "P.N.LEBEDEV" of the Academy of Science of the U.S.S.R. (Fizicheskiy Institut im. P.N.Lebedeva AN SSSR)

PRESENTED BY: SUBMITTED:

16.1.1957

AVAILABLE:

Library of Congress

Card 2/2

SOV/109-3-11-4/13

AUTHORS:

Vitkevich, V.V. and Kokurin, Yu.L.

TITIE:

Measurement of the Phase and Amplitude Fluctuations of the Radio Waves Which Passed Through the Ionosphere (Izmereniye fasovykh i amplitudnykh fluktuatsiy radiovoln, proshedshikh skvoz' ionosferu)

PERIODICAL: Radiotekhnika i Elektronika, 1958, Vol 3, Nr 11, pp 1373 - 1378 (USSR)

ABSTRACT:

For the purpose of this investigation, the ionosphere can be regarded as a screen which produces a certain amount of phase and amplitude modulation. It is therefore possible, by studying the modulation at a distance from the screen, to investigate the behaviour and the structure of the screen itself. This can be done provided the following conditions are fulfilled:

1) The length of the incident wave should be considerably smaller than the dimensions of the screen discontinuities

(at metre waves this condition is always met);

2) The depth of the phase modulation produced by the

screen should be less than 1 radian.

The problem can be studied by employing a double interferometer such as shown in Figure 1. This consists

of two similar entennae, spaced at a distance b.

Inol. Physics in Lebeler AS USSR

Measurement of the Frase and Amplitude Fluctuations of the Radio Waves which Passed Through the Ionosphere

If a plane wave impinges on the antennae at an angle  $\alpha$  the voltage at the input of the receiver  $\lceil \lceil p-1 \rceil$  (Figure 1) is given by Eq (1), where  $\lambda$  is the wavelength,  $\ell_1$  and  $\ell_2$  are the electrical lengths of the cables between the antennae and the receiver, while  $\alpha$  is the base of the interferometer. The voltage at the output of the detector is proportional to the power at the input, as expressed by Eq (2). If the incidence angle  $\alpha$  is varied, the output voltage of the receiver will describe an interference pattern, which can be expressed by Eq (3). If the signal from the antenna system is applied to two receivers which are connected in such a way that the difference in the electrical lengths of their feeder cables is  $\Delta \ell$ , the output voltages of the receiver are given by:

Card 2/6

Measurement of the Phase and Amplitude Fluctuations of the Radio Waves which Passed Through the Tonosphere

$$A_{1} = \frac{1}{2} [1 - \cos 2\pi x],$$

$$A_{2} = \frac{1}{2} [1 - \cos 2\pi (x - x_{1})]$$
(4)

If at the instant corresponding to a phase x<sub>0</sub> the intensity of the signal changes a certain amount, the increments of the receiver output voltages at the points x<sub>0</sub> are given by Eqs (5), where a is the relative change in the signal strength. If both amplitude and phase fluctuations of the stgnal are present, the increments of the output voltages can be expressed by Eqs (7). It is possible to solve Eqs (7) with respect to a and Ax. From Eq (7), it can be seen that the most advantageous arrangement of the interferometer is such in which the phases of the two output voltages differ by x<sub>1</sub>. The above interferometer technique was employed to carry out some

-0ard3/6

Measurement of the Phase and Amplitude Fluctuations of the Radio Waves Which Passed Through the Ionosphere

measurements. The experimental equipment employed two in-phase antennae operating a wavelength of  $\lambda = 3.5 \, \mathrm{m}$ ; the antennae had an area of 59 m² and a beam width of 16° in the horizontal plane and 50° in the meridian? plane. The distance between the antennae was 162.2 m so that the width of the main lobe of the directional pattern was 74'. The antennae were directed towards the radio stars Swan-A and Cassiopea. The receivers were the normal superheterodynes operating at a frequency of 86 Mc/s. The intermediate frequency of each receiver was 9 Mc/s, the overall bandwidth was 0.2 Mc/s and the time constant of the output device was 7 - 9 sec. The results of some preliminary measurements are shown in Figure 2. The continuous curves in the figure correspond to the signal of the source in the absence of fluctuations, while the remaining curves illustrate the fluctuation effects. The curves can be used to determine the amplitude increments  $\triangle A_1$  and  $\triangle A_2$  by the direct measurement of the peaks

Card4/6

Measurement of the Phase and Amplitude Fluctuations of the Radio Waves Which Passed Through the Ionosphers.

in the figure. The values of  $\Delta x$  and a could be evaluated from Eqs (7). The distribution of the quantities  $\Delta \alpha$  and a are shown in Figures 3 and 4, from which it follows that the average value of a is 0.2 and the average value of  $\Delta \alpha$  is 1.6'. The distribution curve of the duration of the fluctuations is given in Figure 5, from which it is seen that the most probable value of the duration is 30 sec. The above values can be used to determine the average phase gradient of the diffracting layer and this has a value of 85 x 10<sup>-5</sup> radians m. Also it is found that the average gradient of the electron density in the ionosphere is 85 x  $10^4$  cm<sup>-2</sup> m<sup>-1</sup>.

There are 5 figures and 8 references, 6 of which are English and 2 Soviet.

Sand 5/6

80V/109-4-1-3/30

AUTHORS: Vitkevich, V. V. and Kokurin, Yu. L.

TITLE: Investigation of the Winds and New-Uniformities in the Ionosphere by Radio-Astronomical Methods (Issledovaniye vetrov i neodnorodnostey v ionosfere radioastronomicheskim metodom)

PERIODICAL: Radiotekhnika i elektronika, 1959, Vol 4, Nr 1, pp 17-20 (USSR)

ABSTRACT: The investigations described were carried out by means of an equipment consisting of 3 antennas and 3 receivers (see Fig.1). The antennas were identical and, in the meridian plane, they had an inclination angle of 5 = 49.5°. Each antenna was in the form of a paraboloidal reflector with a square aperture having an area of 170 m². The width of the directional pattern of the antenna was 21° and the focal distance of the paraboloid was 6.7 m. A half-wave dipole was situated in the focus of each reflector and at a distance of 0.2 \(\lambda\) from it was situated a half-wave reflecting dipole. The 3 receivers were situated in the same place. Each receiver comprised a 2-stage high-frequency amplifier, a heterodyne, a mixer, a 4-stage intermediate-frequency amplifier, a detector, a 2-stage DC push-pull amplifier and a registering device. The intermediate frequency was 10 Mc/s and the band-width of the receiver was 0.4 Mc/s. The anode and heater

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supplies of the receivers were stabilised. The equipment was used to carry cut the measurements on the radio stars in the constellations of Cygnus and Cassiopeia. The measurements were done at a wavelength of 6 m during December 1954 and April 1955. The recordings obtained were in the form shown in Fig.2. From such recordings it was possible to determine the magnitude and the direction of the wind velocity in the ionosphere; the velocity v and the direction  $\beta$  could be evaluated from Eqs.(1), where the meaning of various symbols should be clear by referring to Fig.1. For the West-East velocity component, it was necessary to determine a correction and this was found to be of the order of 20 m/s. The results of the measurements can be summarised as follows: during 13 minutes on 9.12.5+ it was found that the velocity was 70 m/s and  $\beta = 250^\circ$ ; later, the velocity increased rapidly to 90 m/s and  $\beta$  was about 40° for a duration of 6 min; on the 19.12.5+ the velocity was 90 m/s and  $\beta = 280^\circ$ ; these values

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Investigation of the Winds and Non-Uniformities in the Ionosphere by Radic-Astronomical Methods

were constant during 13 minutes; on the 2.4.55 the velocity was 85 m/s and  $\beta = 150^{\circ}$ ; the values were constant during 12 min. The recordings were also used to determine the dimensions of the diffraction spots and it was found that for the direction  $A_1A_2$  the spots had lengths of 1200, 2500 and 2000 m for the above 3 cases, respectively. The paper was read at the Colloquium of the Oscillation Laboratory of the Physics Institute of the Soviet Academy of Sciences on the 8th February, 1956. The paper contains 2 figures and 10 English references.

ASSOCIATION: Fizicheskiy institut im. P. N. Lebedeva, AN SSSR (Physics Institute imeni P. N. Lebedev, Academy of Sciences USSR)

SUBMITTED: May 18, 1957.

Card 3/3

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9,9100 (also 1041, 1046)

Translation from: Referativnyy zhurnal, Geofizika, 1961, No. 2, pp. 42-43,

# 20298 AUTHOR:

Kokurin, Yu. L.

. TITLE: The Shape and Motion of Small Inhomogeneities in the Ionosphere

V sb.: "Dreyfy i neodnorodnosti v donosfere", No. 1, Moscow, AN SSSR, PERIODICAL: 1959, pp. 60-71 (English summary)

Results of the study of the inhomogeneous structure of the ionosphere TEXT: are presented by the spaced reception method with a small base and the use of ... radiosignals from extraterrestrial sources (Signus A, Cassiopeia),  $\lambda = 6m$ . The observations were conducted at Simeiz (44 n. lat., 34 e.long.) from April 1955 to February 1956. The method is briefly described which was applied to the determination of the degree of anisotropy and the drift speed of inhomogeneities by using the correlation analysis. It is shown that the most probable degree of extension of inhomogeneities is 1~1.6 - 1.8, whereat the anisotropy of inhomogeneities in cases of 1 < 5 is not connected with the geomagnetic field, and in case of 1 > 10 the inhomogeneities are extended along the geomagnetic lines of

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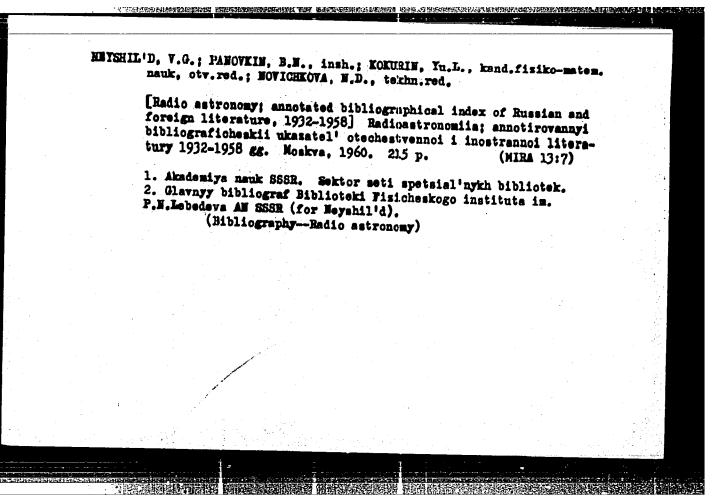
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The Shape and Motion of Small Inhomogeneities in the Ionosphere

force. The most probable dimension of the inhomogeneitles studied is 1 km. No regular seasonal or diurnal variations of this value were observed. It is shown that the life time of inhomogeneities is determined not by diffusion but by another process, most probably by turbulence. Quantitative estimations of the diffusion coefficient and the life time of inhomogeneities are presented. The most probable value of the drift speed was determined to be V = 100 m/sec; no noticeable diurnal, and seasonal variations were observed. The direction of drift changed within wide limits. By night chiefly southward directions could be recorded. The dependence of the state of the ionusphere on the hour of the day was studied. It is shown that the probability of the phenomenon of scintillations of radiostars is highest by night; consequently, the ionosphere is mostly disturbed at this time. In some cases transitions of the ionosphere were observed from the quiet state into the disturbed one, which is explained, in the author's opinion, by the existence of local inhomogeneities in the ionosphere, which represent the regions of decreased electron concentration. The author explains the disagreement of some results of the analysis presented with the data of other investigations by the difference between the latitudes of the observation stations

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S/109/61/006/005/006/027 D201/D303

9,9100

AUTHORS:

Kokurin, Yu.L., Sukhanovskiy, A.N., and Alekseyev, Yu.

I.

TITIE: Investigating of models of large-scale inhomogeneities in the ionosphere using the radioastronomical method

PERIODICAL: Radiotekhnika i elektronika, v. 6, no. 5, 1961, 738 - 746

TEXT: It has already been shown by V.V. Vitkevich, and Yu.L. Kokurin (Ref. 1: Radiotekhnika i elektronika 1957, 2, 7, 826) that the oscillations of the refraction of radiowaves propagated through the whole thickness of the ionosphere are conditioned by the presence in the ionosphere of inhomogeneities with horizontal dimensions of the order of hundreds of kilometers. Again Yu.L. Kokurin (Ref. 2: Radiotekhnika i elektronika 1959, 4, 12, 1985) approximated the evaluation of the dependence of the mean amplitude of oscillations of refraction (R<sub>n</sub>) max on the vertex angle z, and it was

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Investigating of models ...

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shown that this dependence is essentially different for the two models of the layer. In the present article, the authors give the results of measuring the irregular refraction of the ionosphere for two angles simultaneously  $z \approx 0-30^\circ$  and  $z \approx 85^\circ$ . The source of radiation were the sun spots, measurements were made using a vertical naval interferometer in the manner described in Ref. 1 (Op.cit.) The interferometer data were as follows: working frequency f = 207 Mc/s ( $\lambda = 1.45$  m); antenna height over the sea level H = 286.3 m. The results are illustrated of recording the sunrise on December 29, 1958. Further detailed observations were made only during the sunsets, from which basic parameters of large ionosphere inhomogeneities were determined by measuring the periods and amplitudes of refraction oscillations. For each wave (period) of oscillations straight lines, tangential to  $R_{\rm V}^{\rm V}(z)$  at two points were determined at the beginning and end of the period as shown in Fig. 4. The distance between the two points was assumed to be equal to the period of oscillations or to the angular dimension of the inhomogeneity  $\Delta Z$ , and half of the distance of the curve  $R^{\rm V}$  as referred to the

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tangent was taken as the amplitude of the oscillation of refraction (RV) max. Angular dimensions were then transposed into the linear dimension d under the assumption that the distribution of the inhomogeneity was at a height ho = 300 km (Ref. 1: Op.cit.). The value of d oscillation between 100 - 500 km with its most probable value d = 200 - 220 km. The amplitudes of oscillations of refraction (RV) max, averaged over every session of observation, lie basically within the limits 0.5 - 5.0' with the most probable value (RV) max = 2.5 - 3.0'. From the above data the parameters of the two models of inhomogeneities were evaluated as follows: Model 1. Assuming the filmear dimensions d = 200 km its effective thickness of the and the refractive index n = 0.9983 (N = 1.8 · 100 cm-3) the difference between the geometrical and optical thickness of the inhomogeneity; is L = 80 m. Frob Equation (6) obtained by Yu. L. Kokurin (Ref. 2: Radiotekhnika i elektronika, 1959, 4, 12, 1985) the variations of Card 3/76

Investigating of models ...  $\frac{S/109/61/006/005/006/027}{D2C1/D303}$  this difference  $L = \frac{(R_N^V)_{max}}{L} d \frac{\left(1 - \frac{r_0}{r_0 + h_0} \sin s\right)^2 - 2.7 - 3.5 \frac{h^2}{N}}{2\pi}$  (1) (radius of earth -  $r_0$ ) from which  $\frac{\Delta L}{L} = 3.3 - 4.4 \frac{h}{N}$ ; thus if the irregularities in the refraction are due to the presence in the F layer of horizontal gradients, the horizontal changes (with an average period  $\sim 200 \text{ km}$ ) of the optical thickness of large inhomogeneities and of the total number of, electrons in them are  $3.3 - 4.4 \frac{h}{N}$ . Model 2. For the same parameters of inhomogeneities for the wave model the following is obtained using Equation (10) from Yu. L. Ko-kurin (Ref. 2: Op.cit.).  $\frac{\Delta h}{L(2\pi)^3} \frac{r_0}{r_0 + h_0} \sin s^2 \frac{h}{N} = \frac{h}{N} = \frac{h}{N} \frac{h}{N} = \frac{h}{N}$ 

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Investigating of models ...

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It follows that the observed oscillations in the refraction may be attributed to the wave structure of the ionosphere inhomogeneities with a period  $\bar{d}=200$  km and amplitude of the wave  $\Delta h \simeq 0.5$  km. The observations of the irregular refraction near the vertex were carried out in the Crimea (44°N) using a horizontal interferometer consisting of two parabolic antennas spaced in an East-West direction by about  $D=520~\text{m}_1$  the effective beam width was about 150. In order to determine the curves of the dependence of the irregular refraction  $R_n$  on time, the position of the antonna lobes were determined in time units with the origin as the instant of culmination of the source. Observations were made between December 12, 1958 and June 1, 1959 with four cosmic sources. Oraphs are given for every session of observations for  $R_n = f(t)$ . The authors conclude that large-scale ionosphere inhomogeneities represent wave type formations (Model II) with an average horizontal scale (period)  $d \approx 200$  km and the amplitude of the wave  $\Delta h > 0.5$  km. Only an insignificant thickness of the layer seems to have a wave structure. This thickness is <20% of its total effective value. It would Card 5/1/

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Investigating of models ...

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appear as if this part of the inhomogeneity were distributed near the region of maximum electron concentration and has the geometrical thickness < 50 km. There are 5 figures, 1 table and 14 references: 10 Soviet-bloc and 4 non-Soviet-bloc. The references to the English-language publications read as follows: H. Hewish, Proc. Roy. Soc., 1952, 214A, 404; J.P. Wild, J.A. Roberts, J.Atmos and Terr. Physc. 1956, 8, 55; G.N. Munro, Proc. Roy. Soc., 1950, 202, 208; E.N. Bramley, Proc. Roy. Soc., 1953, 220, 39.

ASSOCIATION: Fizicheskiy institut im. P.N. Lebedeva AN SSSR (Institute of Physics im. P.N. Lebedev AS SSSR)

SUBMITTED: June 4, 1960

Card 6/7

ACCESSION NR: AP4019968

\$/0020/64/154/006/1303/1305

AUTHORS: Grasyuk, A.Z.; Zuyev, V.S.; Kokurin, Yu.L.; Kryukov, P.G.;
Kurbaaov, V.V.; Lobanov, V.F.; Moshsherin; V.M.; Sukhanovsky,
A.N.; Cherny\*kh, N.S.; Chuvayev, K.K.

TITLE: Optical moon ranging

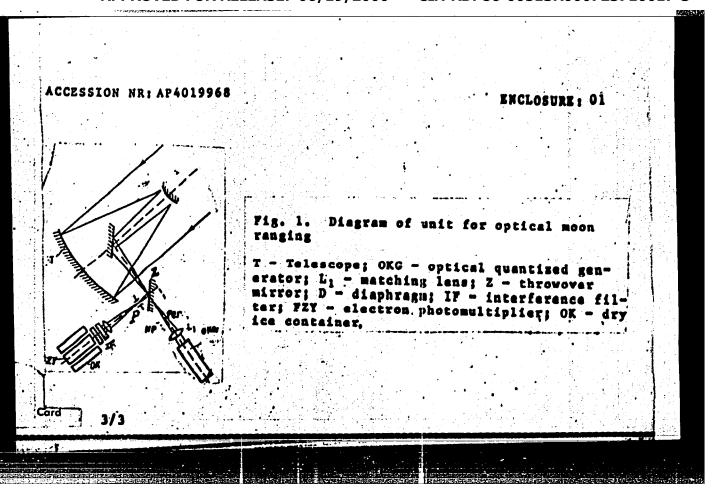
SOURCE: AN SSSR. Doklady\*, v. 154, no. 6, 1964, 1303-1305

TOPIC TAGS: laser, ruby laser, moon ranging, moon
light reflection, celestial ranging, optical ranging

ABSTRACT: The paper describes the preliminary results of moon
ranging with a ruby laser. For the transmission and reception of
the light pulses, a telescope was used with a mirror diameter of 2.6
m. (see Fig. 1 of the Enclosure). The less used was developed by V.S. Zuyev and
P.M. Kryukov and had the following parameters: wavelength 6943A,
pulse energy 50 to 70 joules, pulse duration 2 wasec, diameter
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the light scattering in the atmosphere, the diameter of the spot on
the moon is estimated to be 14 km. For the detection of the signal,

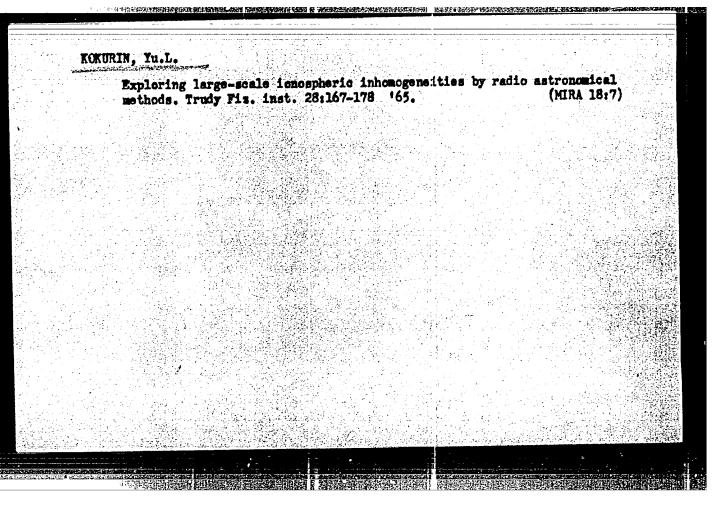
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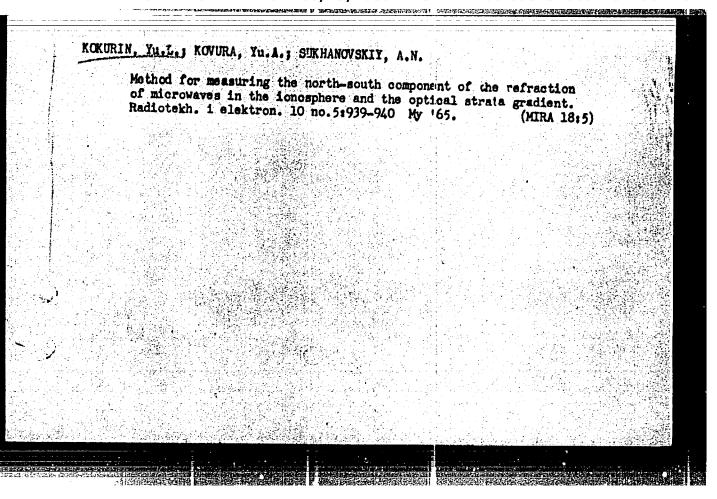
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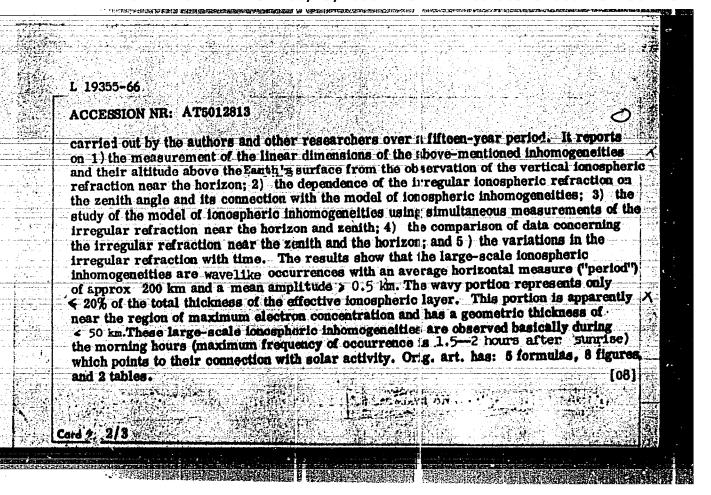
L 52041-65 780/EWT(1)/EWG(\*)/HEC-4/EEG(t)/FCS(k) Pe-5/Pae-2/P1-4/P1-4/ P1-4 VS\_4/GY/VR UR/2504/65/028;000/0129/0134 ACCESSION NR: AT5012810 AUTHOR: Kokurin, Yu. L.; Sorochenko, R. - Radiotelescopes with spherical reflect; rs -<@ SOURCE: AN SESR. Fizicheskiy institut. Trudy, v. 18, 1965. Radioteleskopy (Radio telescopes), 129-134 TOPIC TACS: radictelescope reflector, spherical reflector, radiotelescope reemitter, radiotelescope field ABSTRACT: The construction of large-scale, movable, radiotelescope antennas meets with savere technological difficulties. Consequently, it seemed promising to study radiotelescope systems consisting of a static spherical reflector and a small specially shaped resmitter placed within the tocal region of the mirror (A. K. Head, Nature, 1957, 179, no. 4562). In this paper, submitted in November, 1960 to the enlarged plenery meeting of the Komissi's po radioastronomii (Commission on Radioastronomy), the authors investigate theoretically the shape of the recuitting surface of the spherical radiotelescope and its field of view and ompare its properties with known alternative solutions. A spherical reflector If cover approximately 80% of the sources which an be observed by movable 

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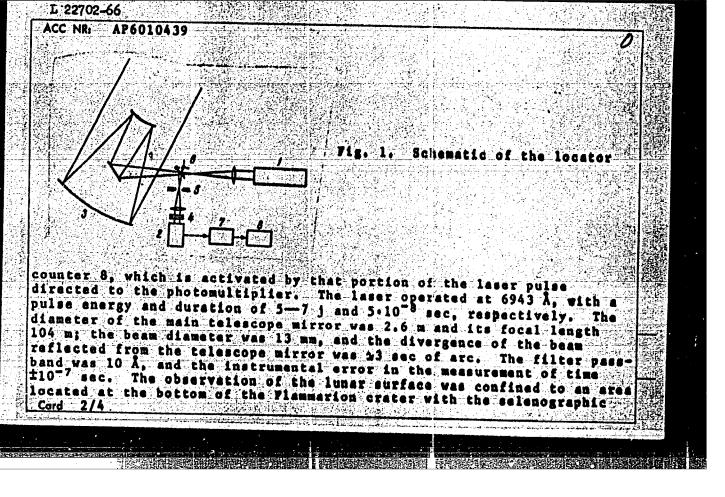
1 19355-66 EWT(1)/FCC/EWA(h)\_\_\_GW UR/2504/65/028/000/0167/017 ACCESSION NR: AT5012818 AUTHOR: Kokurin, Yu. L. TITLE: 14. Studies of large-scale ionospheric inhomogeneities using radio astronosical methods SOURCE: AN SSSR. Fizicheskly institut. Trudy, v. 28, 1965. Radioteleskopy (Radio telescopes), 167-178 TOPIC TAGS: ionospheric inhomogeneity, ionospheric refraction, radio astronomy, electron concentration ABSTRACT: Earlier investigators discovered (see, e. [., W. Rose, E.N. Bramley, Nature, 1949, 164, 355; E.N. Bramley, Proc. Roy. Scc., 1953, 220, 39) that the direction of radio waves from an extraterrestrial source toward an observation point on the earth is subjected to slow irregular changes. The magnitude of these angular oscillations is proportional to the square of the wavelength (B.M. Chikhachev, Radiotekhnika i elektronika, 1960, 5, 9, 1, 359), which supported the assumption that they are caused by irregular radiowave refraction on electron inhomogeneities within the Earth's ionosphere. The present paper gives a theoretical summary of observations

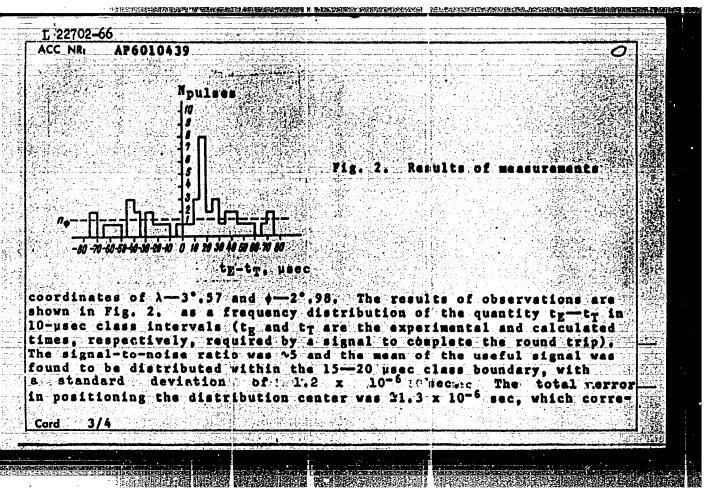


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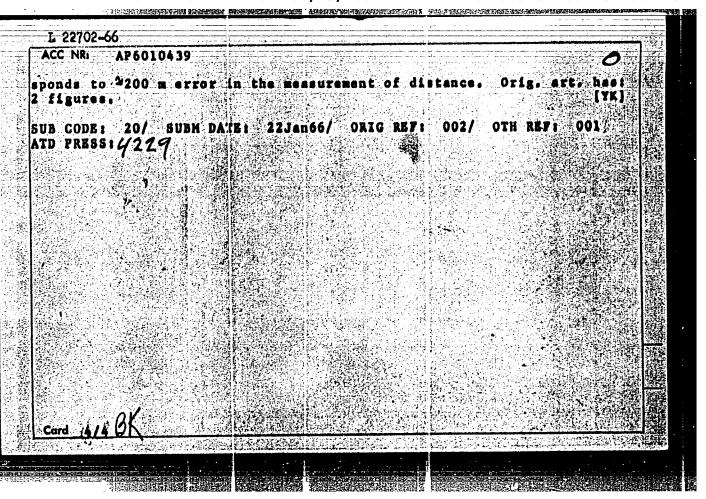
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AUTHOR: Kokurin Yu. V. H.; Sukhanovskiy, A	L.; Kurbasov, V. V.; Lobanov . N.; Chernykh, N. S.	/, V. F.; Hosbsherin,
(Pizicheskiy institut		$\boldsymbol{\beta}$
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SOURCE: Zhurnal eksper v redaktsiyu. Prilosh	imental noy 1 teoretichesko aniye, v. 3, no. 5, 1966, 21	y fiziki. Pie'we
TOPIC TAGS: moon, moo location, optical loca	n_earth_distance,_distance = tion, laser_application	easurement, moon
the distance to the moof the locator is show are fixed rigidly in the ference filter 4 is placed approached by the ference filter 4 is placed	on is given of the experiment on by means of an optical long in Fig. 1. Ruby laser 1 and 1	cator. A schematic and photomultiplier 2
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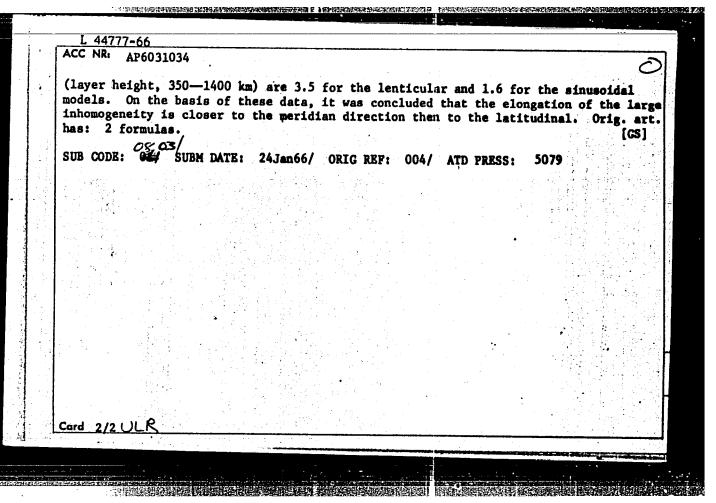


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L 20602-66 DIT(d)/FSS-2/INT(1)/EBC(k)-2/FCC/BNA(h) ACC NR. AP6008279 SOURCE CODE: UR/0109/66/011/003/0439/0444 AUTHOR: Kokurin, Yu. L.; Kovura, Yu. A. ORG: none CIM TITLE: Measuring irregular refraction of radio waves in the ionosphere by means of signals from artificial Earth satellites SOURCE: Radiotekhnika i elektronika, v. 11, no. 3, 1966, 439-444 TOPIC TAGS: electromagnetic wave refraction, ionospheric refraction, artificial satellite ABSTRACT: A theoretical method of isolation of the interference curve free from polarization fading is set forth; information about irregular refraction is obtained by determining the angular position of characteristic points on this curve. The experimental study included reception of a 30-Mc signal from a "Mayak" transmitter borne by the "Elektron-2" satellite; a horizontal half-wave dipole with a reflector and an R-250 superheterodyne receiver were used; the receiver passband was 4-5kc After detection, the signal was applied to a balanced dec amplifier and recorded on Card 1/2 UDC: 621.371.3

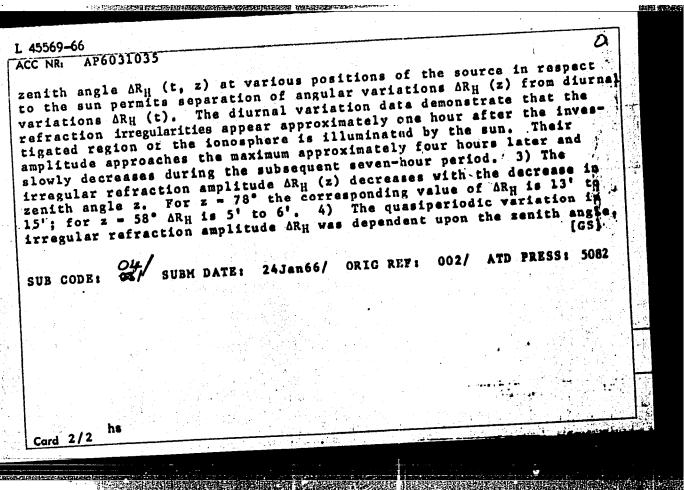
eight was 30 m above s	ent was used in an additional interference level; the antennas of both interference in the state of the state	rometers were so
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分。1972年12月1日 1975年 1 L 44777-66 EWT(1)/FCC GW/WS-2 ACC NR: AP6031034 SOURCE CODE: UR/0109/66/011/009/1687/1688 AUTHOR: Kovura, Yu. A.; Kokurin, Yu. L.; Ovsyankin, M. A. ORG: Physics Institute im. P. N. Lebedev, AN SSSR (Fizicheskiy institut AN SSSR TITLE: Preliminary results in determining the anisometry of large ionospheric inhomogeneities by the radioastronomy, method SOURCE: Radiotekhnika i elektronika, v. 11, no. 9, 1966, 1687-1688 TOPIC TAGS: ionospheric ware propagation, ionospheric inhomogeneity, radu ABSTRACT: To establish the anisometry of large ionospheric inhomogeneities, fluctuations in ionospheric refraction were determined by measuring a 47-Mc emission from the discrete source Cygnus-A within zenith angles of 45°-25°. The study was conducted in the Crimea in November-December 1964. The two interferometers used in the study were oriented at 81° and 204° (clockwise from the south). The error of measuring fluctuations in angles of arrival of the wave did not exceed 1'. A clearly defined anisotropy of incoming wave-angle fluctuations was determined from statistical processing of 19 signal records; the rms values were 5' and 1' for directions 81° and 204°, respectively. The relative dimensions of large ionospheric inhomogeneities and their azimuths were computed for lenticular and sinusoidal models of inhomogeneities. The dimension ratios of large ionospheric inhomogeneities for directions 206° and 84° Card 1/2 UDC: 523.164;621.371



EWT(1)/FCC 45569-66 UR/0109/66/011/009/ ACC, NR. - AP6031035 146 SOURCE CODE: Mont Kokurin, Yu. L.; Oveyankin; ORG: Physics Institute im. P. N. Lebedev. Institut AN SSSR) TITLE: Relating large ionospheric inhomogeneities to the sun according to radioastronomy data 12 SOURCE: Radiotekhnika i elektronika, v. 11, no. 9, 1966, 1688 TOPIC TAGS: ionospheric inhomogeneity, ionospheric propagation ABSTRACT: To determine the dependence of irregular refraction on the zenith angle, the Crimean Scientific Station of the Physics Institute, Academy of Sciences USSR, conducted prolonged radioastronomic observations of two discrete sources, Virgo-A and Cygnus-A, at  $\lambda=6.38$  m in zenith-angle intervals  $z=78^{\circ}-58^{\circ}$  and z=6.38 m in were made using an asymmetrical radio interferometer voriented in an East-West direction. The following results were obtained: 1) An East-West asymmetry of large ionospheric inhomogeneities was noted: fluctustions of radio wave refraction were observed in the Eastern part of the sky but were almost absent in the Western part. 2) An investigation of the dependence of irregular refraction amplitude on the time and the UDC: 523.164:621.371 Card 1/2

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THE PERSON OF TH SOURCE CODE: UR/0293/66/004/003/0414/0426 FACC NR. AF6019595 AUTHOR: Kokurin, Ku. Le.; Kurbasov, V. V.; Lobanov, V. F.; Mozhzhorin, V. M.; Sukhanovskiy, A. N.; Chernykh, N. S. TITLE: On the feasibility of measuring lunar disk and orbital parameters ORG: none SOURCE: Kosmicheskiye issledovaniye, v. 4, no. 3, 1966, 414-426 TOPIC TAGS: lunar albedo, moon, laser application Yu. L. Kokurin and coworkers [1] have reviewed the theoretical ABSTRACT: problems in laser ranging of the moon, with the object of determining more accurate values for several Earth-Moon parameters. The authors discuss methods for 1) obtaining a more detectible reflection signal and 2) using the measured range to compute such parameters as mean lunar orbital radius, lunar disk radius, parallax constant, and Earth equatorial radius. The basic range equation for a reflected electromagnetic signal is taken; as a starting point. The factors are the same as in the radar range equation, except that the return signal varies inversely as the square, rather than as the fourth power, of range, since it is assumed that all the generated laser flux is incident on the Moon. Using an average figure for atmospheric absorption, a lunar albedo of 0.1, and an effective telescope area of 5.3 m2 (actual area of a telescope currently in use), the authors calculate UDC: 523.31.082.5 + 521.61.082.5 Card 1/5

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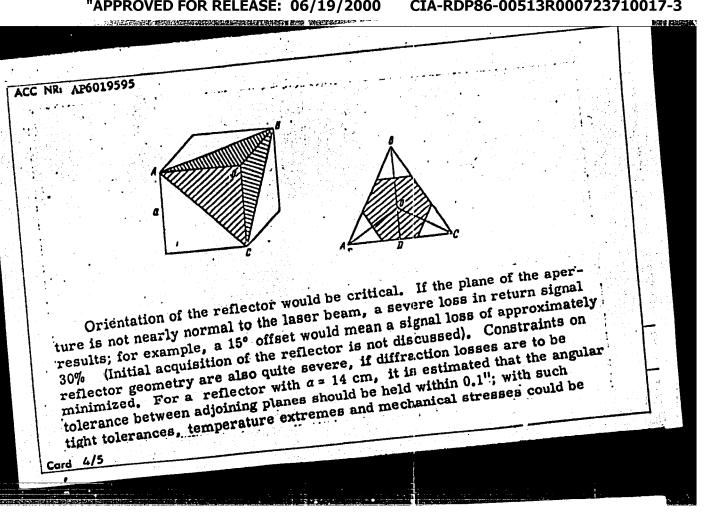
that the relationship between reflected and transmitted energy is

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It follows that with the highest sensitivity photodetectors now available, Wir must be at least 150 joules in order to obtain from the Moon a consistently detectible reflection, i.e., one that does not require statistical analysis to be detected. The pulse must be as short as possible to maximize range resolution; however, present laser pulses of the energy level demanded would have durations of the order of milliseconds, which means a range uncertainty of several hundred kilometers. If Q-switching is used to shorten pulse time, there is an intolerable loss in power amplitude. The conclusion is that only when more powerful short-pulse lasers are developed can there be a significant refinement in lunar ranging measurements.

Factors which degrade the laser technique are also discussed. One of these is the unavoidable divergence of the beam in the atmosphere, estimated at 2" to 3", which would give a lunar spot of some 3.5-5 km across. Contour irregularities within the illuminated area can add to the range uncertainty in the return signal, in the form of range "smear." Owing to the Moon's curvature, a similar effect occurs which increases as a function of

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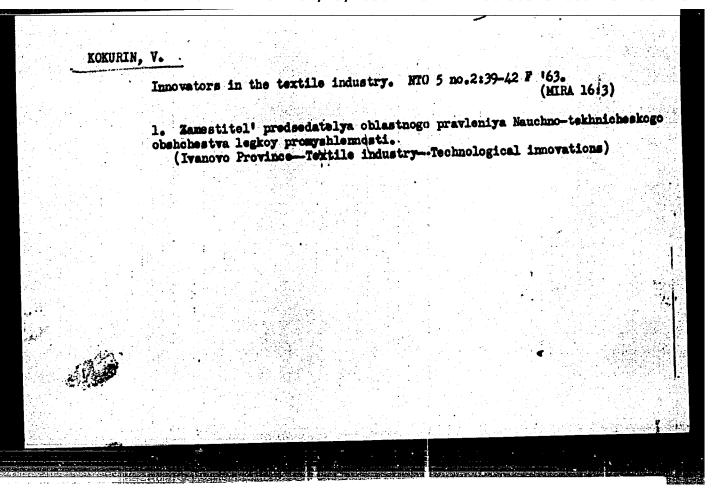
critical factors in reflector performance. Under reasonably good conditions, however, it is calculated that a reflector with  $\beta = 40$  would return an adequate detectible signal to Earth from a Q-switched ruby laser of 4 to 5 joules output.

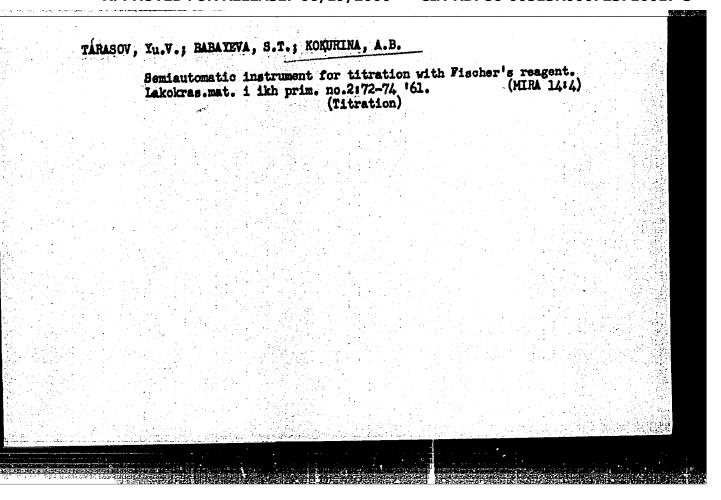
The possibility of confusing a genuine signal with noise or surface rather than reflector return can be minimized by using multiple detection and correlating the results. In fact, if three photomultipliers are used simultaneously, the experiment could be performed in daylight, with a low probability of error.

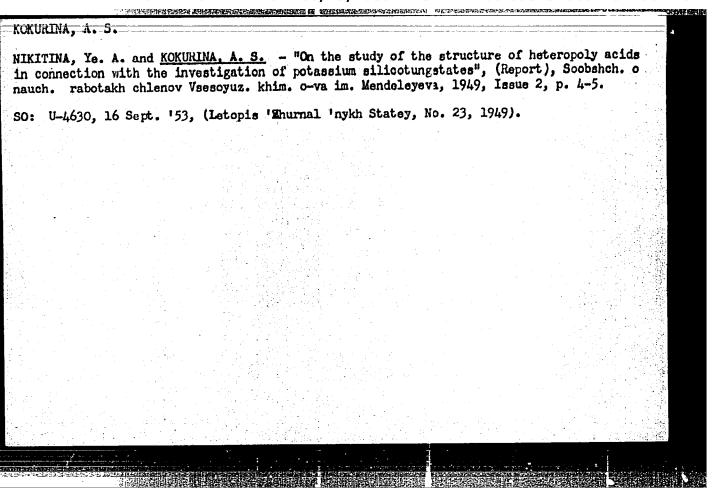
The authors conclude by giving the procedures for calculating mean lunar orbital radius (mean distance between Earth and Moon mass centers), radius of the lunar disk, Earth equatorial radius, and Earth-Moon parallax constant. All of these are obtainable from knowledge of an arbitrary lineof-sight distance from the Earth to the Moon, measured as described above. The calculations show that, with the improved ranging method, parameters such as the Moon's orbital radius and disk radius could be determined to accuracies of several hundreds of meters, a great improvement over the present accuracy of several kilometers. Unfortunately, these accuracy figures do not seem to be tied to any tolerance on the range measurement.

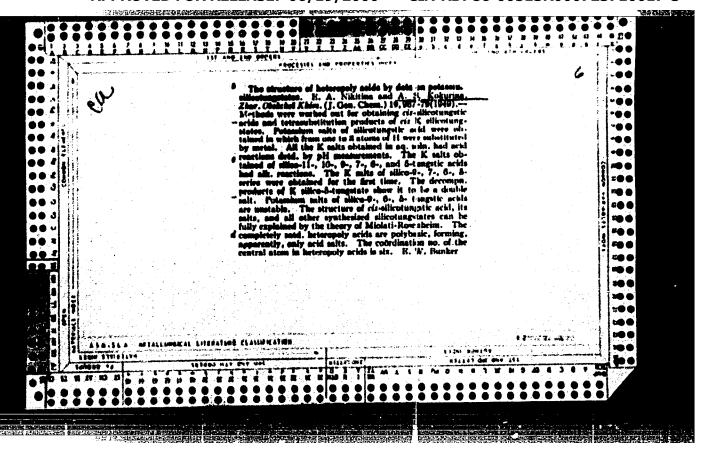
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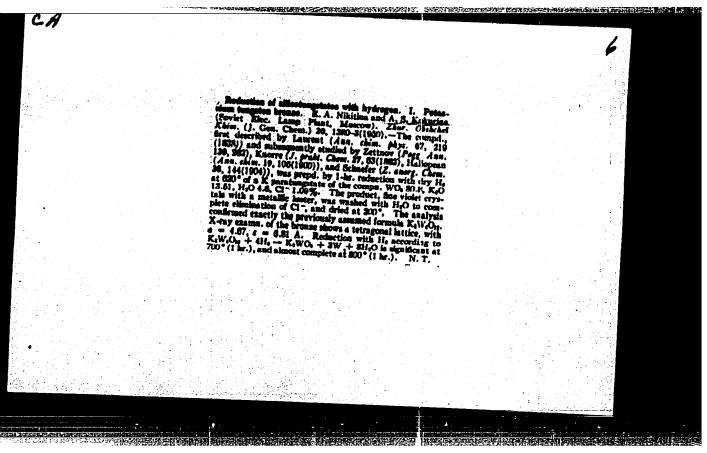
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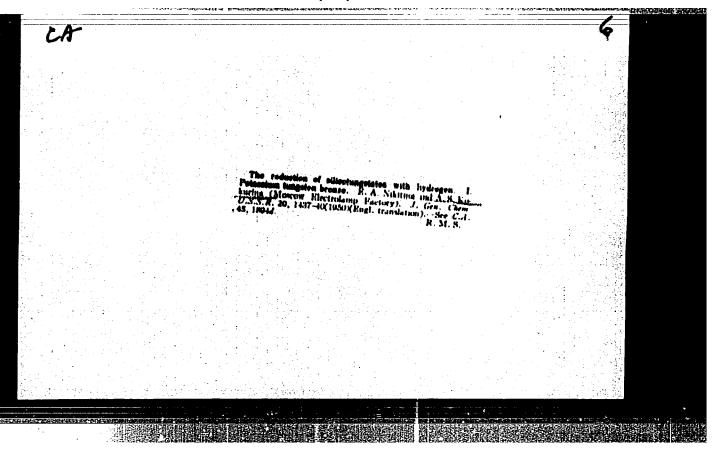


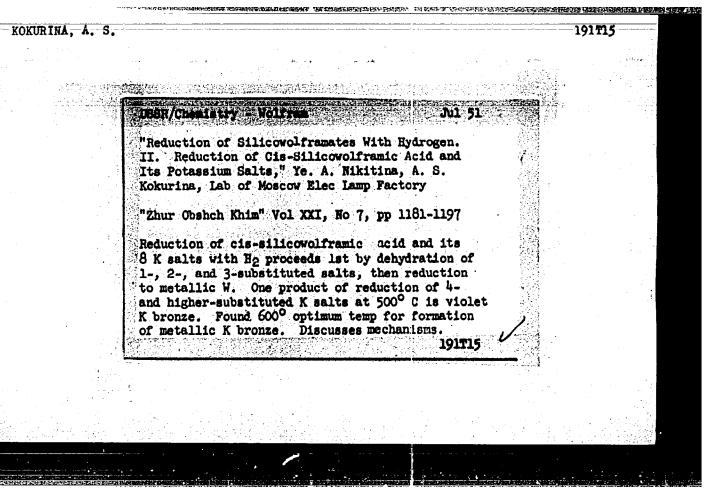




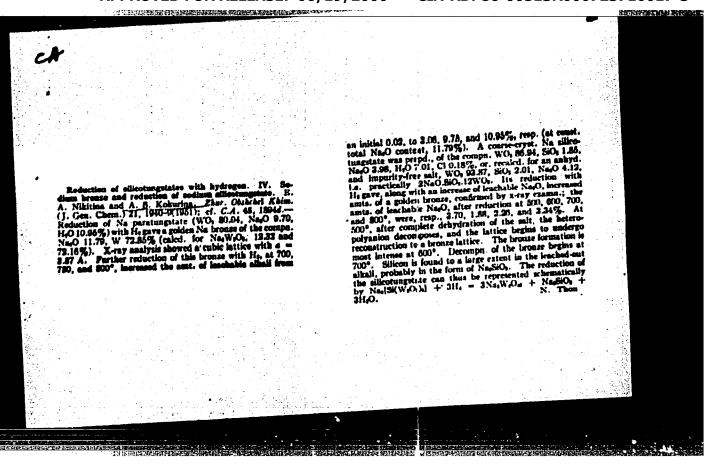








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1. Vtoroy Moskovskiy gosudarstvennyy meditsinskiy institut im. I.V.
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